

Introductory Nuclear, Atomic and Molecular Physics

PHYS-H-405

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Course organization

- Theory:
 - 3 ECTS
 - 1.5 ECTS for nuclear physics (45% of the final note) and 1.5 ECTS for atomic and molecular physics (45% of the final note)
 - Slides for nuclear physics part available on http://metronu.ulb.ac.be/pauly_cours.html
- Exercises:
 - 1 ECTS (0.5 ECTS for each part)
- Laboratories:
 - 1 ECTS
 - Organization: M. Ciccarelli (Maureen.Ciccarelli@ulb.be)
 - 10% of final note → Laboratory reports

References:

- K.S. Krane : Introductory Nuclear Physics (Wiley, 1988)
- K. Heyde : Basic Ideas and Concepts in Nuclear Physics (Institute of Physics, 1994)
- S.S.M. Wong : Introductory Nuclear Physics (Wiley, 1998)
- B.H. Bransden and C.J. Joachain, Physics of Atoms and Molecules (Prentice Hall, 2003)
- B.R. Judd, Operator Techniques in Atomic Spectroscopy (Princeton Landmarks in Physics, 1998)
- W.R. Johnson, Atomic Structure Theory (Lectures on Atomic Physics, Springer, 1998)
- R.D. Cowan, The Theory of Atomic Structure and Spectra (Los Alamos Series in Basic and Applied Sciences, 1981)

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Part I: Nuclear Physics

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2. General properties of nuclei
3. Nuclear models
4. Radioactive decay
5. Alpha decay
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7. Gamma decay
8. Nuclear Fission
9. Nuclear Fusion

Part I: Nuclear Physics

Chapter I: Introduction

Summary

1. Definition
2. Brief history
3. Units and dimensions
4. Structure of nuclei
5. Types of forces
6. Conservation laws
7. Origin of nuclei

Definition

- Nuclear physics = study of atomic nuclei →
 - Theoretical model of atomic nucleus
 - Interaction of particles with nucleus
 - Mechanisms of nuclear reactions
- Applications:
 - Medicine (nuclear medicine: application of radioactive substances in the diagnosis and treatment of disease)
 - Energy production (fission, fusion)
 - Military applications
 - Food-processing (sterilization of food by irradiation)
 - Astrophysics
 - ...

Brief history (1)

1895: Discovery of X-rays by Röntgen

1896: Discovery of radioactivity from uranium by Becquerel

1897: Discovery of the electron by Thomson

1898: Pierre and Marie Curie → other materials are radioactive
→ discovery of Ra and Po

1899: Discovery of α and β rays by Rutherford

1900: Discovery of γ rays by Villard

1903: Discovery of the law of radioactive decay by Rutherford
and Soddy

1905: $E = mc^2$ by Einstein

1908: Discovery of the nucleus by Rutherford

1909: α is a helium nucleus and β is an electron (Rutherford)

Brief history (2)

- 1912: X rays and γ rays are electromagnetic waves (von Laue)
- 1913: Discovery of the notion of isotope (Soddy and Richards)
- 1923: Use of radioactive tracers in biology by von Hevesy
- 1928: Theory of decay based on tunnel effect by Gamow
- 1929: Invention of the cyclotron by Lawrence and Livingston
- 1930: Pauli predicts the existence of the neutrino / Dirac predicts antimatter
- 1932: Discovery of neutron by Chadwick / Discovery of positron by Andersen
- 1934: Fermi theory for β decay
- 1936: Strong force occurs through meson exchange (Yukawa)
- 1936: Lawrence treats leukemia with ^{32}P

Brief history (3)

- 1938: Hahn, Strassman, Meitner and Frisch discover the fission
- 1939: Bethe discovers the nuclear fusion in stars
- 1942: First fission reactor (Fermi)
- 1945: First atomic bomb at Hiroshima
- 1948: Big Bang nucleosynthesis (Alpher, “Bethe”, Gamow)
- 1951: First nuclear reactor producing electricity (EBR-1, Idaho)
- 1952: First hydrogen bomb (Teller, Ulam) / Decision for creation of the CEAN (future SCK-CEN) in Belgium
- 1954: Protontherapy at Berkeley
- 1956: First reactor at critical state in Belgium (BR-1)
- 1961: First PET scan at Brookhaven
- 1964: Gell-Mann and Zweig propose the model of quarks

Brief history (4)

1964: Theory of Brout-Engler-Higgs boson

Mid-1970s: Standard model

1975: First nuclear reactor producing electricity in Belgium
(Doel-1)

1979: Three Mile Island accident (INES 5)

1986: Tchernobyl accident (INES 7)

2011: Fukushima accident (INES 7)

2013: Experimental evidence of BEH boson (CERN)

2020: First protontherapy center in Belgium

2035 (?): First fusion reactor ITER

20??: MIRRHA in Belgium: First accelerator-driven system

20??: SMR in Belgium

Units and dimensions: Typical values

- Size of the atom $\approx 10^{-10}$ m
- Size of the nucleus $\approx 10^{-15}$ m = 1 femtometer (fm) = fermi $\rightarrow$ all nuclei have radius = 2-8 fm
- Typical β or γ decay energy in the range of 1 MeV (megaelectron-volt) = 10^6 eV = $1.6021765 \times 10^{-13}$ J (1 eV = energy gained by a single unit of charge when accelerated through a potential difference of 1 V)
- Unit of mass $\rightarrow$ 1 unified atomic mass unit (u) = $1.6605390 \times 10^{-27}$ kg $\rightarrow$ 1/12 of the mass of an unbound neutral atom of ^{12}C (in ground state and at rest)
- Practically $\rightarrow$ use of mass energy rather than mass $\rightarrow$ multiplication by c^2 ($c = 299\,792\,485\text{ ms}^{-1} \approx 3 \times 10^8\text{ ms}^{-1}$) $\rightarrow 1\text{ u} = 931.502\text{ MeV}$
- Unit of charge $\rightarrow$ elementary charge (e) = $1.6021766209 \times 10^{-19}$ C (proton: e, electron: -e)
- Mean lifetime $\tau = 1/\lambda$ with λ = probability of disintegration per unit time $\rightarrow$ can be 10^{-21} or 10^{19} s

Units and dimensions: Multiplication by a power of c

- Mass $m \rightarrow mc^2$ (energy)
- Momentum $p \rightarrow pc$ (energy)
- Time $t \rightarrow tc$ (length)
- Physical constants:
 - Planck constant: $\hbar = 1.05 \times 10^{-34} \text{ Js} \rightarrow \hbar c = 197.33 \text{ MeVfm}$
 - Proton mass: $m_p = 1.6726 \times 10^{-27} \text{ kg} \rightarrow m_p = 938.27 \text{ MeV}/c^2$
 - Neutron mass: $m_n = 1.6749 \times 10^{-27} \text{ kg} \rightarrow m_n = 939.57 \text{ MeV}/c^2$
 - Electron mass: $m_e = 9.1094 \times 10^{-31} \text{ kg} \rightarrow m_e = 0.511 \text{ MeV}/c^2 \approx m_p/1836$

Structure of nuclei: Nucleons

- Atomic nuclei are quantum bound states of particles called *nucleons*
- Two types of nucleons → positively charged proton and uncharged neutron
- The mass difference between proton and neutron is known with a huge precision: $m_n - m_p = 1.293\,332\text{ MeV}/c^2$
- Nucleons are fermions (spin $\frac{1}{2}$)
- Nucleon is not an elementary particle → an elementary particle has its root mean square (rms) radius = 0

Root mean square radius of proton and neutron (1)

- The rms radius (r_{rms}) of a particle (or charge radius) is defined as the radius of the charge distribution inside the particle →

$$\langle r^2 \rangle_{ch}^{1/2} = \sqrt{\int r^2 \rho_{ch}(\mathbf{r}) d\mathbf{r}}$$

For a proton →

$$\langle r^2 \rangle_{ch,p}^{1/2} \approx 0.87 \text{ fm}$$

with

$$\int \rho_{ch,p}(\mathbf{r}) d\mathbf{r} = 1$$

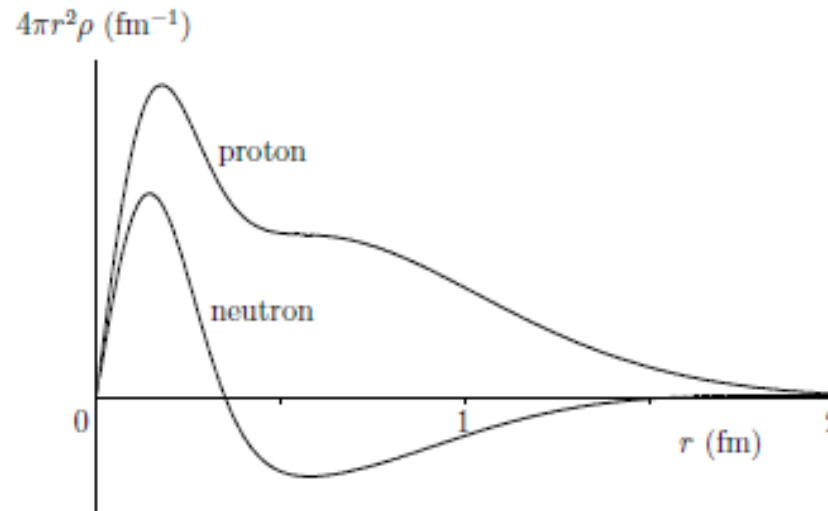
For a neutron →

$$\langle r^2 \rangle_{ch,n} \approx -0.12 \text{ fm}^2$$

with

$$\int \rho_{ch,n}(\mathbf{r}) d\mathbf{r} = 0$$

Root mean square radius of proton and neutron (2)



- The charge density of a nucleon is measured from the analysis of high energy electrons scattered from it
- Practically $\rightarrow$ measurement of charge density is made through their Fourier transform $F(\mathbf{q})$ such as

$$F(\mathbf{q}) = \frac{1}{Z} \int e^{i\mathbf{q}\mathbf{r}} \rho(\mathbf{r}) d\mathbf{r}$$

Structure of nuclei: Magnetic moment

- The spin magnetic moment is the magnetic moment induced by the spin of elementary particles:

$$\mathbf{M} = g \frac{q}{2m} \mathbf{S}$$

with $\mathbf{S}$: the spin, q : the charge, m : the mass, g : the gyromagnetic ratio

- From Dirac theory $\rightarrow$ for charged fermion: $g = 2$, for neutral fermion: $g = 0$ (small corrections from quantum electrodynamics - QED)
- Nucleon is not an elementary particle $\rightarrow g_p = 5.5856947$ and $g_n = -3.826085$

Structure of nuclei: Quarks (1)

- Nucleons are composed of 3 elementary particles: the quarks

Symbol	Spin	Charge	Flavor
u	1/2	+2/3	Up
d	1/2	-1/3	Down
c	1/2	+2/3	Charm
s	1/2	-1/3	Strange
t	1/2	+2/3	Top
b	1/2	-1/3	Bottom

+ anti-quarks

- Quarks have supplementary quantum number: color charge (red, blue, green)
- Anti-quarks have anticolor (antired, antiblue, antigreen)
- Due to the phenomenon of *color confinement* → quarks are never directly observed or found in isolation
- Quarks combine to form hadrons

Structure of nuclei: Quarks (2)

- Theory to be applied to systems of quarks $\rightarrow$ quantum chromodynamics (QCD)
- Combination of 2 quarks (quark + anti-quark) $\rightarrow$ meson (pion, kaon,...)
- Combination of 3 quarks with different colors $\rightarrow$ baryon (proton, neutron, hyperon,...)
- Formation of tetraquarks and pentaquarks seems possible (seems to be observed at the CERN)
- Proton $\rightarrow p = u + u + d$
- Neutron $\rightarrow n = u + d + d$
- Spin of the nucleon ($1/2$) results from the coupling of 3 spins $1/2$

Lepton

- Lepton is an elementary particle
- Spin = $1/2$
- Examples:
 - Electron: charge $-e$, mass $\approx 0.5109989 \text{ MeV}/c^2 \approx 511 \text{ keV}/c^2$
 - Positron: charge $+e$, mass $\approx 511 \text{ keV}/c^2$
 - Muon (« heavy electron »): charge $-e$, $m_\mu \approx 209 m_e$
 - Neutrino: charge 0, mass ≈ 0 but not 0 $\rightarrow m_\nu < 3 \text{ eV}/c^2$, 3 flavors (electron, muon, tau), negative helicity (projection of spin onto the direction of momentum)
 - Antineutrino: same charge and mass than neutrino $\rightarrow$ really different?
 $\rightarrow$ not clear but all experiments have shown positive helicity

Types of forces

Force	Amplitude	Range
Strong nuclear interaction	~ 1	$\sim \text{fm}$
Coulomb interaction (or electromagnetic)	$\sim 1/137$	Infinite
Weak nuclear interaction	$\sim 10^{-5}$	$\sim 10^{-3} \text{ fm}$
Gravitation interaction	$\sim 10^{-39}$	infinite

- Gravitational interaction is negligible compared to the other ones $\rightarrow$ but for systems with a huge number of particles $\rightarrow$ becomes dominating because of the weak total charge of macroscopic systems
- In nuclei $\rightarrow$ Coulomb interaction is not negligible $\rightarrow$ becomes very important when the number of protons $\nearrow$

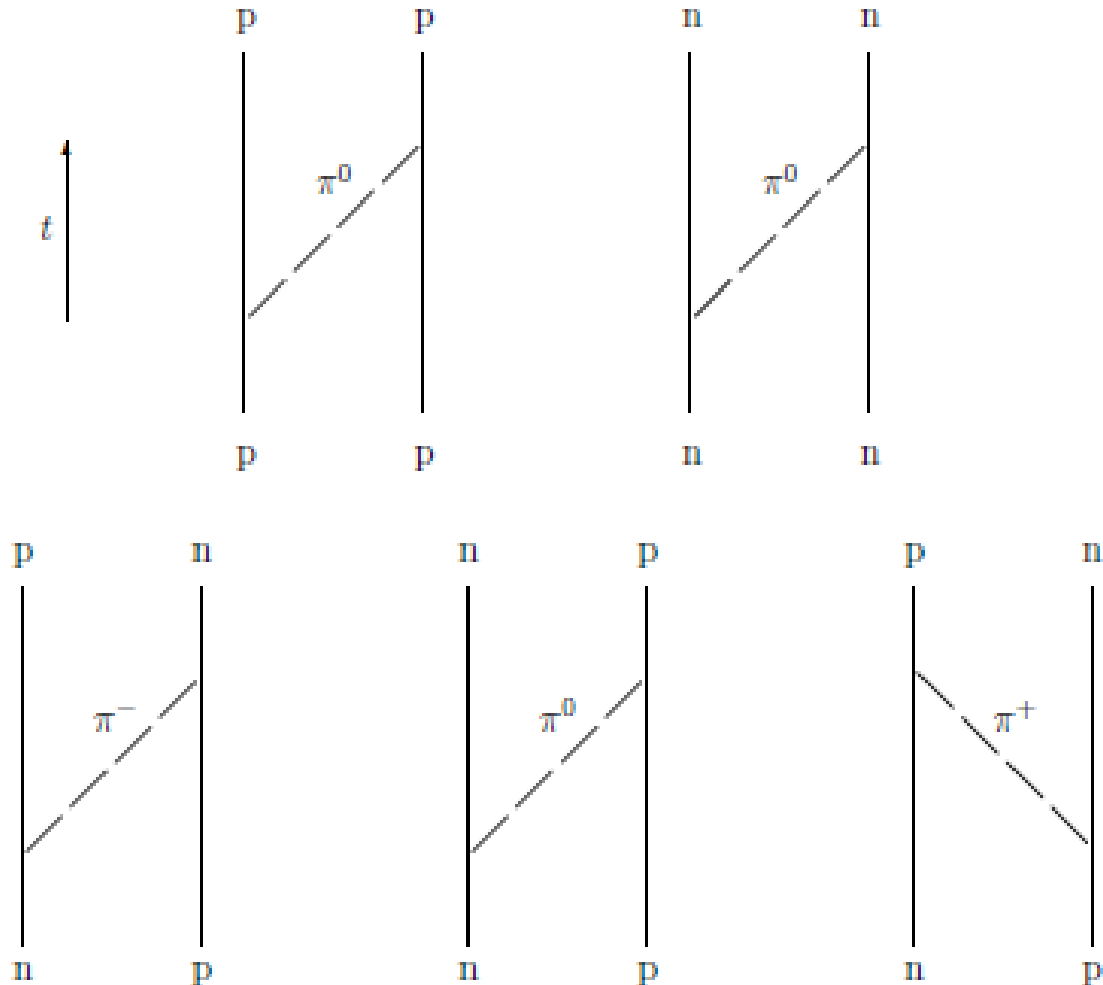
Strong nuclear interaction (1)

- During interactions $\rightarrow$ exchange of virtual particles between particles (Coulomb interaction $\rightarrow$ exchange of photons)
- For strong nuclear interaction between nucleons $\rightarrow$ exchange of mass virtual particles (Yukawa theory) $\rightarrow$ pions (or pi mesons) Π^+ , Π^- and Π^0 (index = charge)
- Due to time-energy uncertainty relation $\rightarrow \Delta t \Delta E \simeq \hbar$
- The energy fluctuation ΔE necessary to have a possible reaction is $\Delta E \sim m_{\Pi} c^2$
- In the time interval Δt the pion can travel a distance $\sim c \Delta t \sim \hbar / (m_{\Pi} c)$ with $\hbar / mc$ the reduced Compton wavelength
- This distance gives the range of the nuclear force

$$m_{\Pi^+} \simeq 139.570 \text{ MeV}/c^2 \quad \Rightarrow \quad \lambda_{\Pi^+} \simeq 1.414 \text{ fm}$$

$$m_{\Pi^0} \simeq 134.977 \text{ MeV}/c^2 \quad \Rightarrow \quad \lambda_{\Pi^0} \simeq 1.462 \text{ fm}$$

Strong nuclear interaction (2)



Feynmann diagrams for n/p interactions

Strong nuclear interaction (3)

- Same physical mechanism for interaction between 2 protons, 2 neutrons and 1 proton/1 neutron $\rightarrow$ exchange of same type particle $\rightarrow$ very similar interactions $\rightarrow$ property of *charge independence*
- The strength of the strong interaction between any pair of nucleons is the same independently of the nucleon type (protons or neutrons)
- Charge independence is not perfect because π^0 is necessary exchanged for nn and pp interactions and np interactions can be done with various π $\rightarrow$ as potentials are not exactly the same $\rightarrow$ not perfect charge independence

Weak nuclear interaction (1)

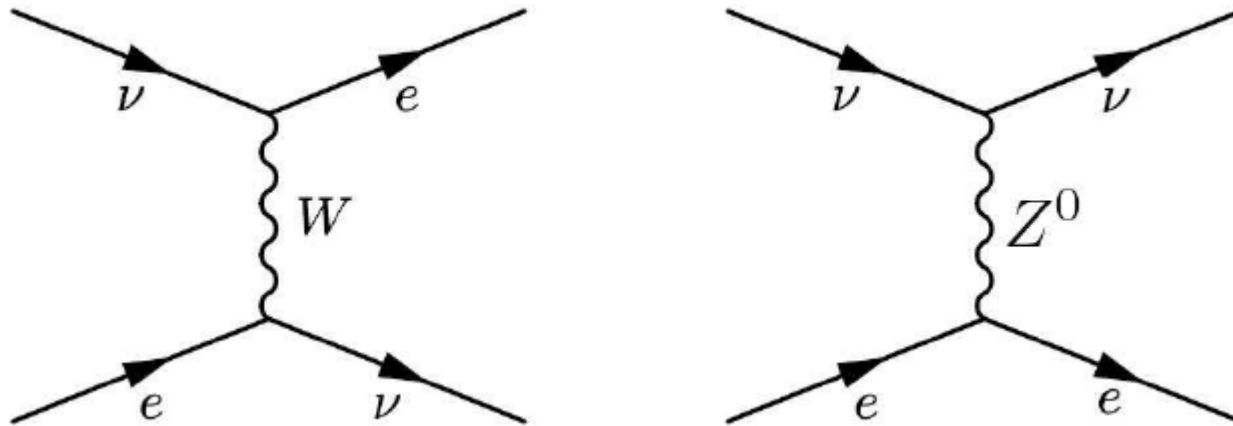
- Weak nuclear interaction always neglected in all calculations of nuclear structure
- This effect only appears in processes forbidden to strong nuclear interactions $\rightarrow \beta$ disintegration
- Yukawa theory can be adapted to weak interaction $\rightarrow$ weak gauge bosons (W^+ , W^- , Z^0) mediate the weak interactions

$$m_W \simeq 80.4 \text{ GeV}/c^2 \quad \Rightarrow \quad \lambda_W \simeq 0.00246 \text{ fm}$$

$$m_Z \simeq 91.19 \text{ GeV}/c^2 \quad \Rightarrow \quad \lambda_Z \simeq 0.00217 \text{ fm}$$

- Short range $\rightarrow$ weakness of the interaction

Weak nuclear interaction (2)



- The electroweak theory unifies weak interaction and electromagnetic interaction
- This theory shows that at high energy ($E > m_Z c^2$) $\rightarrow$ the 2 interactions have the same order of magnitude

Conservation laws (1)

- A *conservation law* states that a particular physical quantity does not changed (is conserved) during a physical process
- A conservation law is *exact* if it was never contradicted
- A conservation law is *approximated* if the considered quantity is conserved in certain classes of physics processes but not in all (e.g. conservation of parity)
- Principals conservation laws are:
 - conservation of energy and momentum (exact)
 - conservation of angular momentum (exact)
 - conservation of electric charge (exact)
 - conservation of baryon number (approximated)
 - conservation of lepton number (approximated)
 - ...

Conservation laws (2)

- Conservation of baryon number $\rightarrow$ conservation of the total number of nucleons (neutrons and protons) minus the total number of anti-nucleons (antiprotons and antineutrons) $\rightarrow$
 $p + d \nrightarrow p + p + \gamma$ is not allowed (charge YES, baryon number No)
 $\rightarrow \gamma + d \rightarrow p + n$ is allowed
- Conservation of electronic lepton number (L_e) $\rightarrow$ conservation of the difference between the total number of [electrons and electron neutrinos] and the total number of [positrons and electronic antineutrinos] $\rightarrow$

$$L_e \equiv N(e^-) + N(\nu_e) - N(e^+) - N(\bar{\nu}_e)$$

$\rightarrow \bar{\nu}_e + n \nrightarrow e^- + p$ is NOT allowed but $\rightarrow \bar{\nu}_e + n \rightarrow e^+ + n$ YES

- Remark 1: it exists two other types of charged leptons $\rightarrow \mu^\pm$ and $\tau^\pm \rightarrow$ similar conservation laws with L_μ and L_τ

Conservation laws (3)

- Remark 2: some recent experiments on neutrino oscillations shows that the only truly conserved number is the sum of the 3 lepton numbers: $L = L_e + L_\mu + L_\tau$
- Remark 3: $p \nrightarrow e^+ + \gamma$ is not possible (conservation of baryon and lepton number) but a « possible » theory predicts the disintegration of proton (with a characteristic lifetime $\tau_p \approx 10^{29}$ years $\approx 10^{19}$ times the age of Universe)
- Remark 4: neutron is instable $\rightarrow n \rightarrow p + e^- + \bar{\nu}_e$ (with $Q = (m_n - m_p - m_e)^2 \approx 0.782$ MeV and $\tau_n \approx 885.7$ s)

Conservation laws: Stability of nuclei

- Conservation of energy $\rightarrow$ 1 particle of mass m and charge q can spontaneously decay into i particles only if (with Q the liberated energy):

$$m > \sum_i m_i \text{ or equivalently } \rightarrow Q = mc^2 - \sum_i m_i c^2 > 0$$

- Conservation of charge $\rightarrow q = \sum_i q_i$
- These laws implies that electron and positron are stable (no particle of same charge but of smaller mass)

Origin of nuclei

- Nuclei in nature were built by nuclear reactions since the « Big Bang »
- Begin not well known $\rightarrow$ system at very high $T \rightarrow$ free quarks and gluons (bosons mediating between quarks) = quark-gluon plasma
- Time $t \nearrow \rightarrow T \searrow \rightarrow$ quarks and gluons combine to form hadrons and nucleons $\rightarrow$ proton $\rightarrow$ neutron and neutron $\rightarrow$ proton $\rightarrow$ but as $m_p < m_n \rightarrow$ system with smallest mass is favored
- $t \nearrow$ more $\rightarrow T \searrow \rightarrow$ collisions $\rightarrow$ apparition of bounded systems (with $A \leq 7$) $\rightarrow$ fixed situation = 87% of protons and 13% neutrons
- Abundance: 74% H + 23-25% He ($4p \rightarrow {}^4\text{He} + 2e^+ + 2\nu_e + n\gamma$) + ...
- Other nuclei are formed during star explosion: supernova $\rightarrow$ a lot of nuclei are unstable and decay into other nuclei $\rightarrow$ process of formation + decay is called *nucleosynthesis*