Chapter II: Quantum physics

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Introduction (1)

- Quantum mechanics is a fundamental theory in physics → describes nature and explain properties of microscopic molecules, atoms and subatomic particles
- Small objects have characteristics of both particles and waves
- Work of Planck in 1900 (analyzing blackbody radiation) and Einstein in 1905 (analyzing the photoelectric effect) → light energy is delivered not smoothly and continuously as a wave but instead in concentrated bundles or « quanta »
- In 1924 de Broglie postulated that associated with a « particle » moving with momentum p is a « wave » of wavelength λ = h/p (h = Planck's constant) → λ = de Broglie wavelength

Introduction (2)

- Probabilistic character of quantum mechanics → results of measurements are given as a collection of possibilities → each of them associated to a given probability
- The goal of quantum physics is to explain and foresee the evolution in time of a physical system (collection of particles)
- Principles of quantum mechanics are expressed in a series of postulates
- Remarks : in general cases → kinetic energy of particles is much smaller than rest energy → nonrelativistic quantum mechanics

Postulates (1)

First postulate:

At each instant the state of a physical system is represented by a ket $|\psi\rangle$ in the space of states

Comments:

- The space of states is a vector space
- − The space of states includes the concept of inner product \rightarrow the inner product associates a complex number to any two states \rightarrow

$$(|\psi\rangle, |\phi\rangle) \equiv \langle \psi | \phi \rangle = \int dx \psi^*(x) \phi(x)$$

- A function f(x) can be evaluated in the $|\psi\rangle$ system \rightarrow expectation value of $f \rightarrow \langle \psi | f | \psi \rangle = \int dx \psi^*(x) f(x) \psi(x)$

Postulates (2)

Second postulate:

Every observable attribute of a physical system is described by an operator that acts on the kets that describe the system

Comments:

− By convention an operator A acting on a ket $|\psi>$ is denoted by left multiplication →

$$A: |\psi\rangle \to |\psi'\rangle = A |\psi\rangle$$

− In the context of wave-mechanics \rightarrow state is replaced by wavefunction $\psi(\mathbf{x}) \rightarrow \text{example of momentum operator (1D) } p = -i\hbar d/dx \rightarrow$

$$p |\psi\rangle = p\psi(x) = -i\hbar \frac{d\psi(x)}{dx}$$

Postulates (3)

Third postulate:

The only possible result of the measurement of an observable A is one of the eigenvalues of the corresponding operator A

Comments:

- Origin of the word « quantum »
- For every operator → there are special states that are not changed by the action of an operator (except for being multiplied by a constant) → they are the eigenstates and the constant numbers are the eigenvalues of the operator:

$$A \left| \psi_a \right\rangle = a \left| \psi_a \right\rangle$$

Postulates (4)

Fourth postulate:

When a measurement of an observable A is made on a state $|\psi\rangle$ the probability of obtaining an eigenvalue a_n is given by the square of the inner product of $|\psi\rangle$ with the eigenstate $|a_n\rangle$ $\rightarrow |\langle a_n |\psi\rangle|^2$

Comments:

- − The states are (normally) assumed to be normalized to unity → $|<\psi|\psi>| = 1$ and $|<a_j|a_k>| = \delta_{jk}$
- The complex number $|\langle a_n | \psi \rangle|$ is known as the « probability amplitude » or « amplitude » to measure a_n as the value for A in the state $|\psi\rangle$

Postulates (5)

Fifth postulate:

Immediately after the measurement of an observable A has yielded a value a_n , the state of the system is the normalized eigenstate $|a_n|$

Comments:

- This is known as the « collapse of the wavepacket »

Postulates (6)

Sixth postulate:

The state $|\psi(t)\rangle$ of each non-relativistic quantum system is a solution of the Schrödinger equation depending on the time \rightarrow

$$i\hbar \frac{d}{dt} |\psi(t, \boldsymbol{r})\rangle = H |\psi(t, \boldsymbol{r})\rangle$$

Comments:

− *H* is the Hamiltonian operator \rightarrow it represents the total energy of the particle of mass *m* in the potential field *V* \rightarrow time-independent or stationary Schrödinger equation \rightarrow

$$H\psi(\mathbf{r}) = E\psi(\mathbf{r})$$
 with
 $H = -\frac{\hbar^2}{2m}\Delta\Psi + V(\mathbf{r})$

Problems in one dimension: Free particle (1)

• No forces $\rightarrow V(x) = 0 \rightarrow$



- First term exp \rightarrow wave travelling in the positive *x* direction
- Second term exp \rightarrow wave travelling in the negative x direction

Problems in one dimension: Free particle (2)

- Intensities of the waves are given by $|A|^2$ and $|B|^2$
- No boundary conditions \rightarrow no restrictions on the energy $E \rightarrow$ all values of *E* give solutions to the equation
- No convergence at +∞ or -∞ → normalization condition cannot be applied in this case
- Different normalization system \rightarrow source such as an accelerator located at $x = -\infty$ emitting particles at a rate *I* (particles s⁻¹) with momentum $p = \hbar k$ in the positive *x* direction \rightarrow particles traveling in the positive *x* direction $\rightarrow B = 0 \rightarrow$ particle current $j = (\hbar k/m)|A|^2$

Problems in one dimension: Step potential $E > V_0(1)$

$$V(x) = 0 \quad x < 0, \text{ region } 1$$

= $V_0 \quad x > 0, \text{ region } 2$ with $V_0 > 0$

- For $x < 0 \rightarrow$ as free particle with $k = k_1 = (2mE/\hbar)^{1/2}$
- For $x > 0 \rightarrow$ $\psi_2(x) = Ce^{ik_2x} + De^{-ik_2x}$ $k_2^2 = \frac{2m(E - V_0)}{\hbar^2}$
- Boundary conditions give $\rightarrow A+B = C+D$ and $k_1(A-B) = k_2(C-D)$
- Particle comes from source at x = -∞ → A term represents the incident wave / B term is the reflected wave / C term is the transmitted wave / D = 0 (no possibility of reflection after the step)

Problems in one dimension: Step potential $E > V_0$ (2)

• Reflection coefficient or probability \rightarrow

$$R = \frac{|B|^2}{|A|^2} = \left(\frac{1 - k_2/k_1}{1 + k_2/k_1}\right)^2$$

• Transmission coefficient

$$T = \frac{k_2}{k_1} \frac{|C|^2}{|A|^2} = \frac{4k_2/k_1}{(1+k_2/k_1)^2}$$

• Application in nucleon-nucleon scattering problems



Image taken from: K.S. Krane, Introductory Nuclear Physics, Wiley, Oboken, 1988 Problems in one dimension: Step potential $E < V_0$

• In region 2 \rightarrow

 $\psi_2(x) = Ce^{k_2x} + De^{-k_2x} \qquad k_2^2 = \frac{2m(V_0 - E)}{\hbar^2}$

- As first term $\rightarrow \infty$ for $x \rightarrow \infty \rightarrow C = 0$
- Important difference compared to classical mechanics → All classical particles are reflected at the boundary ↔ the quantum mechanical wave packet can penetrate a short distance into the forbidden region



Image taken from: K.S. Krane, Introductory Nuclear Physics, Wiley, Oboken, 1988 Problems in one dimension: Barrier potential $E > V_0(1)$

$$V(x) = 0 \quad x < 0, \text{ region } 1$$
$$= V_0 \quad 0 \ge x \ge a, \text{ region } 2$$
$$= 0 \quad x > a, \text{ region } 3$$

Solutions in regions 1, 2 and 3 \rightarrow

$$\psi_{1}(x) = Ae^{ik_{1}x} + Be^{-ik_{1}x}$$

$$\psi_{2}(x) = Ce^{ik_{2}x} + De^{-ik_{2}x}$$

$$\psi_{3}(x) = Fe^{ik_{3}x} + Ge^{-ik_{3}x}$$

$$k_{1}^{2} = k_{3}^{2} = \frac{2mE}{\hbar^{2}} \qquad k_{2}^{2} = \frac{2m(E - V_{0})}{\hbar^{2}}$$

Problems in one dimension: Barrier potential $E > V_o(2)$

- We use continuity conditions at *x* = 0 and at *x* = *a*
- We assume particles coming from $x = -\infty \rightarrow G = 0$
- After calculations → transmission coefficient becomes



Problems in one dimension: Barrier potential $E < V_o(1)$

• Expressions in region 1 and 3 are identical and in region 2 \rightarrow

$$\psi_2(x) = Ce^{k_2 x} + De^{-k_2 x} \qquad \qquad k_2^2 = \frac{2m(V_0 - E)}{\hbar^2}$$

- As x varies between 0 and a in region $2 \rightarrow C$ and $D \neq 0$
- After calculations → transmission coefficient becomes

$$T = \left(1 + \frac{1}{4} \frac{V_0^2}{E(V_0 - E)} \sinh^2 k_2 a\right)^{-1}$$

Classically *T* = 0 → the particle is not permitted to enter the forbidden region (negative kinetic energy) ↔ the quantum wave can penetrate the barrier → nonzero probability to find the particle beyond the barrier.

Problems in one dimension: Barrier potential $E > V_o(2)$

- This phenomenon is called barrier penetration or quantum mechanical tunneling or tunnel effect
- Important applications in nuclear physics $\rightarrow \alpha$ decay and fission



Image taken from: K.S. Krane, Introductory Nuclear Physics, Wiley, Oboken, 1988





Image taken from: K.S. Krane, Introductory Nuclear Physics, Wiley, Oboken, 1988



- The walls are absolutely impenetrable → particle is trapped between x = 0 and x = a
- Inside the solution of the Schrödinger equation is $\Rightarrow \psi(x) = A \sin kx + B \cos kx$
- Continuity condition at $x = 0 \rightarrow \psi(0) = 0 \rightarrow$ true only for B = 0
- At $x = a \rightarrow$ the continuity condition gives \rightarrow

$$A\sin kx = 0$$

1D problem: Infinite well (2)

- As A = 0 is not acceptable $\rightarrow \sin ka = 0$ $\rightarrow ka = n\pi$ for n = 1,2,3,...
- Quantification condition on energy \rightarrow

$$E_n = \frac{\hbar^2 k^2}{2m} = \frac{\hbar^2 \pi^2}{2ma^2} n^2$$

 Corresponding states ψ are bound states (potential confines particle in a certain region of space) →

$$\psi_n = \sqrt{\frac{2}{a}} \sin \frac{n\pi x}{a}$$



Image taken from: K.S. Krane, Introductory Nuclear Physics, Wiley, Oboken, 1988

1D problem: Finite well (1)

$$V(x) = V_0 |x| > a/2$$

= 0 |x| < a/2

• Bound-state solutions (with $E < V_0$) are

$$\psi_{1}(x) = Ae^{k_{1}x} + Be^{-k_{1}x} \quad \text{for } x < -a/2$$

$$\psi_{2}(x) = Ce^{ik_{2}x} + De^{-ik_{2}x} \quad \text{for } -a/2 \le x \le a/2$$

$$\psi_{3}(x) = Fe^{k_{1}x} + Ge^{-k_{1}x} \quad \text{for } x > a/2$$

$$k_1^2 = \frac{2m(V_0 - E)}{\hbar^2} \qquad k_2^2 = \frac{2mE}{\hbar^2}$$

1D problem: Finite well (2)

- Wave function has to be finite in region $1 \rightarrow \text{for } x = -\infty \rightarrow B = 0$
- Wave function has to be finite in region $3 \rightarrow \text{for } x = \infty \rightarrow F = 0$
- Due to continuity at $x = \pm a/2 \rightarrow$

$$k_2 \tan \frac{k_2 a}{2} = k_1$$
 $-k_2 \cot \frac{k_2 a}{2} = k_1$

 Eq. cannot be resolved directly → numerical or graphical methods → graphical solutions easiest with eq. in the form →

1D problem: Finite well (3)

- Right side \rightarrow circle of radius $P \leftrightarrow$ left side \rightarrow tangent function
- Solutions are given by the points where the circle intersects the tangent
- Therefore the number of solutions is determined by the radius $P \rightarrow$ by the depth V_0 of the well (for infinite well \rightarrow infinite number of bound states)
- For P < π/2: only one bound state ↔ for π/2 < P < π: two bound states
- Technique that allows to estimate the depth of the nuclear potential → for deuteron: only one bound state

1D problem: Finite well (4)



Solutions for P = 6 (as an example) \rightarrow 4 solutions at α = 1.345, 2.679, 3.985, 5.226

Images taken from: K.S. Krane, Introductory Nuclear Physics, Wiley, Oboken, 1988 1D problem: Simple Harmonic Oscillator (1)

$$V(x) = \frac{1}{2}kx^2$$

 Any reasonably well-behaved potential can be expanded in a Taylor series about the point x₀ →

$$V(x) = V(x_0) + \left(\frac{dV}{dx}\right)_{x=x_0} (x - x_0) + \frac{1}{2} \left(\frac{d^2V}{dx^2}\right)_{x=x_0} (x - x_0)^2 + \dots$$

- If x₀ is a potential minimum → the second term = 0 and since the first term is a constant → the interesting term is the third term
- Near its minimum the system behaves like a simple harmonic oscillator → The study of the simple harmonic oscillator is important for a large number of systems

1D problem: Simple Harmonic Oscillator (2)

• To solve the Schrödinger equation \rightarrow change of $\psi \rightarrow$

$$\psi(x) = h(x) \exp\left(-\alpha^2 x^2/2\right)$$

- h(x) is a simple polynomial function in x and $\alpha^2 = (km)^{1/2}/\hbar$
- The degree of the polynomial → the highest power of x that appears is determined by the quantum number n that labels the energy states →

$$E_n = \hbar \omega_0 \left(n + \frac{1}{2} \right)$$
 with $n = 0, 1, 2, ...$

• where $\omega_0 = (k/m)^{1/2}$ is the classical angular frequency of the oscillator

1D problem: Simple Harmonic Oscillator (3)

	n	E_{κ}	$\psi_n(x)$		
	0	$\frac{1}{2}\hbar\omega_0$	$\pi^{-1/4} e^{-\alpha^2 \pi^3/2}$		
	1	$\frac{3}{2}\hbar\omega_0$	$2^{-1/2}\pi^{-1/4}(2\alpha x) e^{-\alpha^2 x^2/2}$		
mages taken from: K.S. Krane, <i>ntroductory Nuclear</i> Physics, Wiley, Dboken, 1988	2	$\frac{5}{2}\hbar\omega_0$	$2^{-3/2}\pi^{-1/4}(4\alpha^2 x^2 - 2) e^{-\alpha^2 x^2/2}$		
	3	$\frac{7}{2}h\omega_0$	$(1/4\sqrt{3}\pi^{1/4})(8\alpha^3x^3-12\alpha x)e^{-\alpha^2x^2/2}$		
	4	$rac{9}{2}\hbar\omega_0$	$(1/8\sqrt{6} \pi^{1/4})(16\alpha^4 x^4 - 48\alpha^2 x^2 + 12) e^{-\alpha^2 x^2/2}$		
	$E_n = \hbar \omega_0 (n + \frac{1}{2})$				
		$\psi_n(x) = (2^n n! \sqrt{2^n n!})$	$(\pi^{-1/2}H_{\nu}(\alpha x) e^{-a^2x^2/2})$		
		where $H_{\alpha}(\alpha x)$ is a Hermite polynomial			



- Results for the probabilities resemble those of finite well
- Energy levels are equally spaced
- Potential is infinitely deep → number of bound states is infinite

3D problem: Infinite Cartesian Well (1)

 $\begin{array}{rcl} V(x) & = & \infty & x < 0, \, x > a, \, y < 0, \, y > a, \, z < 0, \, z > a \\ & = & 0 & 0 \le x \le a, \, 0 \le y \le a, \, 0 \le z \le a \end{array}$

- The particle is confined to a cubic box of dimension *a*
- Inside the well \rightarrow the Schrödinger equation is

$$-\frac{\hbar^2}{2m}\left(\frac{\partial^2\psi}{\partial x^2} + \frac{\partial^2\psi}{\partial y^2} + \frac{\partial^2\psi}{\partial z^2}\right) = E\psi(x, y, z)$$

• With solutions \rightarrow

$$\psi_{n_x,n_y,n_z}(x,y,z) = \sqrt{\frac{8}{a^3}} \sin \frac{n_x \pi x}{a} \sin \frac{n_y \pi y}{a} \sin \frac{n_z \pi z}{a}$$
$$E_{n_x,n_y,n_z} = \frac{\hbar^2 \pi^2}{2ma^2} \left(n_x^2 + n_y^2 + n_z^2\right)$$

3D problem: Infinite Cartesian Well (2)

- *n_x*, *n_y*, *n_z* are independent integers > 0
- The lowest state (ground state) has quantum numbers (n_x, n_y, n_z) = (1,1,1)
- The probability distribution has a maximum at the center of the box falling gradually to 0 at the walls like sin²



Image taken from: K.S. Krane, Introductory Nuclear Physics, Wiley, Oboken, 1988

3D problem: Infinite Cartesian Well (3)

- The first excited state has 3 possible sets of quantum numbers: (2,1,1), (1,2,1) and (1,1,2)
- Each of these distinct and independent states has a different wave function → a different probability density → different expectation values of the physical observables
- Attention → they have the same energy → situation called degeneracy
- Degeneracy is extremely important for atomic structure → how many electrons can be in each atomic subshell

3D problem: Infinite Spherical Well (1)

- In general \rightarrow use of spherical coordinates with potential depending only on *r* (not θ or ϕ)
- Separable solutions of the form $\psi(r,\theta,\phi) = R(r)\Theta(\theta)\Phi(\phi)$
- The central potential V(r) appears only in the radial part of the equation $R(r) \rightarrow$ the angular parts can be solved directly
- Differential equation for $\Phi(\phi)$ (with m_l^2 the separation constant) \rightarrow

$$\frac{d^2\Phi}{d\phi^2} + m_l^2\Phi = 0$$

• The solution is

$$\Phi_{m_l}(\phi) = \frac{1}{\sqrt{2\pi}} \exp\left(im_l\phi\right)$$

• The m_{l} = 0, \pm 1 \pm 2,...

3D problem: Infinite Spherical Well (2)

• The equation for $\Theta(\theta) \rightarrow$

$$\frac{1}{\sin\theta} \frac{d}{d\theta} \left(\sin\theta \frac{d\Theta}{d\theta} \right) + \left[l(l+1) - \frac{m_l^2}{\sin^2\theta} \right] \Theta = 0$$

- Where *I* = 0,1,2,3... and m_I = 0, \pm 1 \pm 2,...
- The solutions $\Theta_{lm_l}(\theta)$ are expressed as a polynomial of degree l in $\sin\theta$ or $\cos\theta$
- Together and normalized, $\Phi_{m_l}(\phi)$ and $\Theta_{lm_l}(\theta)$ give the spherical harmonics $Y_{lm_l}(\theta, \phi)$
- These functions give the angular part of the solution to the Schrödinger equation for any central potential *V*(*r*)
- For example → give the spatial properties of atomic orbitals responsible for molecular bonds

3D problem: Infinite Spherical Well (3)

ł	m_{ℓ}	$Y_{\ell m_\ell}(\theta,\phi) = \Theta_{\ell m_\ell}(\theta) \Phi_{m_\ell}(\phi)$
0	0	$(1/4\pi)^{1/2}$
1	0	$(3/4\pi)^{1/2}\cos\theta$
1	± 1	$\mp (3/8\pi)^{1/2}\sin\theta e^{\pm i\phi}$
2	0	$(5/16\pi)^{1/2}(3\cos^2\theta - 1)$
2	± 1	$\mp (15/8\pi)^{1/2} \sin\theta \cos\theta \ e^{\pm/4}$
2	± 2	$(15/32\pi)^{1/2}\sin^2\theta \ e^{\pm 2i\phi}$
	$\Phi_{m_{\ell}}(\phi) = \frac{1}{\sqrt{2\pi}} e^{im_{\ell}\phi}$	
	$\Theta_{\ell m_{\ell}}(\theta) = \left[\frac{2\ell+1}{2}\frac{(\ell-m_{\ell})!}{(\ell+m_{\ell})!}\right]^{1/2}$	$P_{\ell}^{m_{\ell}}(\theta)$
	where $P_{\ell}^{m_{\ell}}(\boldsymbol{\theta})$ is the a	ssociated Legendre polynomial

Image taken from: K.S. Krane, Introductory Nuclear Physics, Wiley, Oboken, 1988

3D problem: Infinite Spherical Well (4)

• For a given $V(r) \rightarrow$ radial equation is

$$-\frac{\hbar^2}{2m}\left(\frac{d^2R}{dr^2} + \frac{2}{r}\frac{dR}{dr}\right) + \left[V(r) + \frac{l(l+1)\hbar^2}{2mr^2}\right]R = ER$$

- The *I*(*I*+1) term is an addition to the potential → « centrifugal potential » acts like a potential that keeps the particle away from the origin when *I* > 0
- Example of infinite spherical well \rightarrow

$$V(x) = \infty \quad r > a$$
$$= 0 \quad r < a$$

Inside the well → solution are expressed in terms of the spherical Bessel functions j_l(kr)

3D problem: Infinite Spherical Well (5)

- Continuity condition at *r* = *a* to find the energy eigenvalues → *j_l(ka)* = 0
- Transcendental equation → numerical solutions → tables of the spherical Bessel functions to find the zeros for any given value of *l*
- For example → *l* = 0 → from the tables *j*₀(*x*) = 0 at *x* = 3.14, 6.28, 9.42, 12.57,...
- For *l* = 1 → first few zeros of *j*₁(*x*) at x = 4.49, 7.73, 10.90, 14.07

$j_0(kr) = \frac{\sin kr}{kr}$
$j_1(kr) = \frac{\sin kr}{\left(kr\right)^2} - \frac{\cos kr}{kr}$
$j_{2}(kr) = \frac{3\sin kr}{(kr)^{3}} - \frac{3\cos kr}{(kr)^{2}} - \frac{\sin kr}{kr}$
$j_{\ell}(kr) \cong \frac{(kr)^{\ell}}{1 \cdot 3 \cdot 5 \cdots (2\ell + 1)} \qquad kr \to 0$
$j_{\ell}(kr) \cong \frac{\sin(kr - \ell\pi/2)}{kr} \qquad kr \to \infty$
$j_{\ell}(kr) = \left(-\frac{r}{k}\right)^{\prime} \left(\frac{1}{r}\frac{d}{dr}\right)^{\prime} j_{0}(kr)$

Image taken from: K.S. Krane, Introductory Nuclear Physics, Wiley, Oboken, 1988

3D problem: Infinite Spherical Well (6)

- Since $E = \hbar^2 k^2 / 2m \rightarrow$ solutions for the allowed values of *E* energies
- Repeating process for *I* = 2, *I* = 3,... → construction of the spectrum of the energy states



Image taken from: K.S. Krane, Introductory Nuclear Physics, Wiley, Oboken, 1988

3D problem: Infinite Spherical Well (7)

- The levels are also degenerate → since *E* depends only on *l* → the wave functions with different *m_l* values all have the same energy
- Since m_l is restricted to the values 0, ± 1, ± 2, ..., ±l → there are exactly 2l + 1 possible Y_{lml} for a given l → each level has a degeneracy of 2l + 1

3D problem: Simple Harmonic Oscillator (1)

- Central oscillator potential $\rightarrow V(r)=1/2kr^2$
- For all central potentials \rightarrow the angular part of the solution to the Schrodinger equation is $Y_{Im_I}(\theta, \phi) \rightarrow$ we have only to find the solution to the radial equation
- As in 1D → the solution is the product of an exponential and a finite polynomial
- The energy levels are given by

$$E_n = \hbar \omega_0 \left(n + \frac{3}{2} \right)$$
 with $n = 0, 1, 2, ...$

3D problem: Simple Harmonic Oscillator (2)

- No dependence on / for the energy but not all / values are permitted
- From the mathematical solution of the radial equation → / can be at most equal to n and takes only even or only odd values as n is even or odd
- Example \rightarrow For n = 5, the permitted values of l are 1, 3, and 5
- Since *E* do not depend on *m_l* → additional degeneracy of 2*l* + 1 for each *l*
- For *n* = 5 level has a degeneracy of [(2 * 1 + 1) + (2 * 3 + 1) +
- (2 * 5 + 1)] = 21
- Degeneracy equals to 1/2(n+1)(n+2)

3D problem: Simple Harmonic Oscillator (3)

п	l	En	<i>R</i> (<i>r</i>)
0	0	$\frac{3}{2}\hbar\omega_0$	$(2\alpha^{3/2}/\pi^{1/4}) e^{-\alpha^2 r^2/2}$
1	1	$\frac{5}{2}h\omega_0$	$(2\alpha^{3/2}\sqrt{2}/\sqrt{3}\pi^{1/4})(\alpha r) e^{-\alpha^2r^2/2}$
2	0	$\frac{2}{2}\hbar\omega_0$	$(2\alpha^{3/2}\sqrt{2}/\sqrt{3}\pi^{1/4})(\frac{3}{2}-\alpha^2 r^2) e^{-\alpha^2 r^2/2}$
2	2	$\frac{2}{2}\hbar\omega_0$	$(4\alpha^{3/2}/\sqrt{15}\pi^{1/4})(\alpha^2 r^2) e^{-\alpha^2 r^2/2}$
3	1	$\frac{2}{2}\hbar\omega_0$	$(4\alpha^{3/2}/\sqrt{15}\pi^{1/4})(\frac{5}{2}\alpha r - \alpha^3 r^3) e^{-\alpha^2 r^2/2}$
3	3	$\frac{9}{2}h\omega_0$	$(4\alpha^{3/2}\sqrt{2}/\sqrt{105}\pi^{1/4})(\alpha^3 r^3) e^{-\alpha^2 r^2/2}$
4	0	$\frac{11}{2}\hbar\omega_0$	$(4\alpha^{3/2}\sqrt{2}/\sqrt{15}\pi^{1/4})(\frac{15}{8}-\frac{5}{2}\alpha^2r^2+\frac{1}{2}\alpha^4r^4)e^{-\alpha^2r^2/2}$
4	2	$\frac{11}{2}\hbar\omega_0$	$(4 \alpha^{3/2} \sqrt{2} / \sqrt{105} \pi^{1/4}) (\frac{7}{2} \alpha^2 r^2 - \alpha^4 r^4) e^{-\alpha^2 r^2/2}$
4	4	$\frac{11}{2}\hbar\omega_{0}$	$(8\alpha^{3/2}/3\sqrt{105}\pi^{1/4})\alpha^4r^4e^{-\alpha^2r^2/2}$



Images taken from: K.S. Krane, Introductory Nuclear Physics, Wiley, Oboken, 1988

Commutator

• Mathematical definition of Commutator \rightarrow

[A,B] = AB - BA

- This is equal to 0 if they commute and something else if they do not commute
- It is known that you cannot know the value of two physical values at the same time if they do not commute (Heisenberg's principle)
- If we can find a set of operators commuting with H → this set is complete in the sense that all energy eigenstates are uniquely labelled by the eigenvalues of the above operators
- A key property of central potential problems is that the angular momentum operators commute with the Hamiltonian

(Orbital) Angular momentum (1)

- In solutions of the 3D Schrödinger equation → prominent role of the quantum number l
- In atomic physics → label the different electron wave functions and give information about their spatial behaviour
- This angular momentum quantum number has the same function in all 3D problems involving central potentials *V*(*r*)
- In classical physics $\rightarrow I = r \times p$
- In quantum mechanics → components of *p* have to be replaced by their equivalent operators:

$$p_i = -i\hbar \frac{\partial}{\partial x}$$

(Orbital) Angular momentum (2)

- Result for which a central potential is considered \rightarrow having a wave function $R(r)Y_{Im_{I}}(\theta, \phi) \rightarrow$ independent of the form of R(r)
- It is simple to calculate $l^2 \rightarrow$

 $\langle \boldsymbol{l}^2 \rangle = \hbar l(l+1)$

- As in classical physics for central potentials → the angular momentum is a constant of the motion
- The atomic substates with a given / value are labeled using spectroscopic notation → same spectroscopic notation in nuclear physics: s for / = 0, p for / = 1, ...

l value	0	1	2	3	4	5	6
Symbol	S	р	d	f	g	h	i

Image taken from: K.S. Krane,

Introductory Nuclear Physics,

Wiley, Oboken, 1988

(Orbital) Angular momentum (3)

- Direction of *I* → barrier imposed by the uncertainty principle in quantum mechanics → it is permitted to know exactly only one component of *I* at a time → when one component is known the other two components are completely indeterminate
- By convention $\rightarrow z$ component is chosen \rightarrow

$$\langle l_z \rangle = \hbar m_l$$

with m_l = 0, \pm 1 \pm 2,...

Spin angular momentum

- Spin angular momentum = Intrinsic angular momentum = Spin
- Nucleons (like electrons) have spin quantum number $s = \frac{1}{2}$
- Usual calculations for angular momenta \rightarrow

$$\langle s^2 \rangle = \hbar s(s+1)$$

 $\langle s_z \rangle = \hbar m_s \quad (m_s = \pm \frac{1}{2})$

• Useful to imagine the spin as a vector **s** with possible *z* components = $\pm \frac{1}{\hbar}$

Total angular momentum (1)

The total angular momentum *j* combines both the spin and orbital angular momentum of a particle or system →

$$j = l + s$$

• Behaviour of total angular momentum \rightarrow

$$\langle j^2 \rangle = \hbar j (j+1)$$

 $\langle j_z \rangle = \langle j l_z + s_z \rangle = \hbar m_j$

We have m_j = -j, -j + 1,..., j - 1, j (j is the total angular momentum quantum number)

Total angular momentum (2)

- As $m_s = \pm \frac{1}{2} \rightarrow m_j = m_1 + m_s = m_1 \pm \frac{1}{2}$
- m_i is always an integer $\rightarrow m_j$ is half-integral $\rightarrow j$ is half-integral \rightarrow

$$j = l + \frac{1}{2}$$
 or $j = l - \frac{1}{2}$

- The *j* value is noted as a subscript in spectroscopic notation
- For *I* = 1 (p states) → two possible *j* values: *I* + 1/2 = 3/2 and *I* 1/2 = 1/2
- These states are written as $p_{3/2}$ and $p_{1/2}$

General addition of angular momenta

- General rules for the addition of two angular momenta J_1 and $J_2 \rightarrow J = J_1 + J_2$
- When j₁ and j₂ (corresponding quantum numbers) are "added" the maximal and minimal values of j are j_{max} = j₁ + j₂ and j_{min} = |j₁ - j₂|
- The allowed *j*-values in this interval are j = $|j_1 j_2|$, $|j_1 j_2|$ + 1,..., $j_1 + j_2$
- If both j₁ and j₂ are half-integral or if both are integers → the possible quantum numbers j are integers ↔ in the opposite case the resulting j-values become half-integral
- Properties known as triangular inequality and noted \rightarrow

$$|j_1 - j_2| \le j \le j_1 + j_2$$

Parity (1)

- The parity operation causes a reflection of all of the coordinates through the origin: $r \rightarrow -r$
- In Cartesian coordinates $\rightarrow x \rightarrow -x / y \rightarrow -y / z \rightarrow -z$
- In spherical coordinates \rightarrow r \rightarrow r / $\theta \rightarrow \pi$ θ / $\phi \rightarrow \phi$ + π
- If a system is left unchanged by the parity operation → none of the properties should change as a result of the reflection
- The values of the observable quantities depend on $|\psi|^{\,\rm 2}$ \rightarrow

If
$$V(\mathbf{r}) = V(-\mathbf{r}) \to |\psi(\mathbf{r})|^2 = |\psi(-\mathbf{r})|^2$$

Parity (2)

- Consequence $\rightarrow \psi(-\mathbf{r}) = \pm \psi(\mathbf{r})$
- In case $\psi(-\mathbf{r}) = +\psi(\mathbf{r}) \rightarrow \text{positive or even parity}$
- In case $\psi(-\mathbf{r}) = -\psi(\mathbf{r}) \rightarrow$ negative or odd parity
- If the potential V(*r*) is unchanged by the parity operation → the resulting wave functions must have either even or odd parity (mixed-parity wave functions are not permitted)
- For finite well \rightarrow potential symmetric with respect to the parity operation: $V(x) = V(-x) \rightarrow$ even or odd parity for the solutions



Parity (3)

• In 3D the parity operation applied to $Y_{lml}(\theta, \phi)$ gives \rightarrow

$$Y_{lm_l}(\pi - \theta, \phi + \pi) = (-1)^l Y_{lm_l}(\theta, \phi)$$

 Central potentials depend only on the magnitude of *r* are thus invariant with respect to parity → their wave functions have odd arity if *l* is odd and even parity if *l* is even

Parity (4)

- The wave function for a system of many particles = the product of the wave functions for the individual particles → the parity of the combined wave function is even if the combined wave function represents any number of even-parity particles or an even number of odd-parity particles ↔ it is odd if there is an odd number of odd-parity particles
- Thus nuclear states can be assigned a definite parity (odd or even) → indicated along with the total angular momentum for that state → example:

$$\frac{3}{2}^{+}$$
 or $\frac{5}{2}^{-}$

Fermi's golden rule

- Fermi's golden rule is a formula that describes the transition rate (probability of transition per unit time) from one energy eigenstate of a quantum system into other energy eigenstates effected by a weak perturbation
- We consider the system to begin in an eigenstate $|i\rangle$ of a given Hamiltonian H_0
- We consider the effect of a (possibly time-dependent) perturbing Hamiltonian H'.
- The transition probability per unit of time λ from the state $|i\rangle$ to a set of final states $|f\rangle$ is given to first order in the perturbation by

$$\lambda = \frac{2\pi}{\hbar} |\langle f| H' |i\rangle|^2 \rho$$

ρ is the density of final states (number of continuum states per unit of energy)